
NASA/ADS

Ionospheric F-region plasma density topography maps using OI 777.4 nm all-sky images

Show affiliations

Abalde, J.; de Freitas, T.; Fagundes, P.; Sahai, Y.; Pillat, V.; Becker-Guedes, F.; Soubhia, B.

Observations of the OI 777.4 nm nightglow emission using an all-sky imaging system and ionospheric sounding using a Canadian Advanced Digital Ionosonde CADI have been recently carried out at São José dos Campos 23°21' S 45°86' W and Brasília 22°53' S 45°57' W respectively. The OI 777.4 nm emission comes mainly from the radiative recombination process O_e to O^+O followed by radiative process O^+5P to O^+5S $h\nu_{\lambda 777.4\text{ nm}}$. It is known that to a good approximation the peak electron density is proportional to $J_{777.4}^{1/2}$ where $J_{777.4}$ is the vertical column emission rate. Using the near simultaneous optical imaging and ionospheric sounding observations on four selected geomagnetically quiet nights without any presence of large-scale plasma depletions we have obtained the linear correlation coefficient β between the F-region peak electron density and OI 777.4 nm emission intensity. Using β and OI 777.4 nm emission images topography maps of F-region plasma density have been obtained for several nights both with and without plasma bubbles. The topography maps are then used to study the dynamics of the equatorial ionospheric anomaly as well as plasma depleted regions. Some of the topographic maps show the equatorial ionospheric anomaly region having high f_oF_2 values of the order of 15-17 MHz with the surrounding areas having much lower f_oF_2 values in the range of 8-12 MHz. Also the presence of plasma bubbles in the topographic maps on some nights.

Publication:


36th COSPAR Scientific Assembly. Held 16 - 23 July 2006, in Beijing, China. Meeting abstract from the CDROM, #1321

Pub Date:

2006

Bibcode:

2006cosp...36.1321A

 Feedback/Corrections? (</feedback/correctabstract?bibcode=2006cosp...36.1321A>)